Radioheliograph

Berni Alder

The Culgoora Radioheliograph John Paul Wild, 1967

The Culgoora Radioheliograph CSIRO. Division of Radiophysics, Commonwealth Scientific and Industrial Research Organization (Australia). Division of Radiophysics, 1968* Golden Years of Australian Radio Astronomy Wayne Orchiston, Peter

Robertson, Woodruff T. Sullivan III, 2022-01-20 The evolution of Australian radio astronomy from 1945 to 1960 has been studied in detail by numerous historians of science in recent years. This Open Access book is the first to present an overview of this remarkable chapter in Australian science. The book begins in the post-war period, as the Radiophysics Laboratory in Sydney switched from secret wartime research on radar to peacetime applications of this new technology. Next follows the detection of radio waves from space and the ensuing transformation of this fledgling science into the dominant research program at the Radiophysics Lab. Drawing from this history, the book shows how by 1960 the Radiophysics Lab had become the largest and most successful radio astronomy group in the world. The final chapter presents an overview of Australian radio astronomy from 1960 to the present day, as Australia prepares to co-host the multi-national, multi-billiondollar Square Kilometre Array. Nearly 300 high-quality images complement the text, drawn from a wide range of sources including the extensive collection held by the CSIRO Radio Astronomy Image Archive. The book will be an essential reference for readers interested in the scientific and cultural development of radio astronomy. This book is published open access under a CC BY 4.0 license.

Solar and Space Weather Radio Physics Bin Chen, Dale E. Gary, Nicole

Vilmer, 2021-07-28

Radio Astronomy Berni Alder, 2012-12-02 Methods in Computational Physics, Volume 14: Radio Astronomy is devoted to the role of the digital computer both as a control device and as a calculator in addressing problems related to galactic radio noise. This volume contains four chapters and begins with a technical description of the hardware and the special data-handling problems of using radioheliography, with an emphasis on a selection of observational results obtained with the Culgoora radioheliograph and their significance to solar physics and to astrophysics in general. The subsequent chapter examines interstellar dispersion, its influence on time resolution, methods for its measurement, and removing its effect. This chapter also outlines resolution and sampling problems, as well as the computation of the average pulse profile. This chapter surveys also the intensity variations over time scales from microseconds to millions of years and over radio frequency ranges from 40 MHz to 10 GHz. Another chapter highlights the special case of rotational aperture synthesis and its problems, followed by a presentation of data processing necessary to convert the parameters observed with an aperture synthesis telescope into an intensity distribution of part of the sky. The last chapter discusses some data-acquisition and data reduction techniques, as well as some selected problems in data interpretation in spectral-line radio astronomy. This book will be of great importance to geoscientists, physicists, and mathematicians.

The Culgoora Radioheliograph ,1970

Four Pillars of Radio Astronomy: Mills, Christiansen, Wild, Bracewell R.H. Frater, W.M. Goss, H.W. Wendt, 2017-10-28 This is the story of Bernie Mills, Chris

Christiansen, Paul Wild and Ron Bracewell, members of a team of radio astronomers that would lead Australia, and the world, into this new field of research. Each of the four is remembered for his remarkable work: Mills for the development the cross type instrument that now bears his name; Christiansen for the application of rotational synthesis techniques; Wild for the masterful joining of observations and theory to elicit the nature of the solar atmosphere; Bracewell for his contribution to imaging theory. As well, these Four Pillars are remembered for creating a remarkable environment for scientific discovery and for influencing the careers of future generations. Their pursuit of basic science helped pave the way for technological developments in areas ranging from Wi-Fi to sonar to medical imaging to air navigation, and for underpinning the foundations of modern cosmology and astrophysics.

Introduction to Solar Radio Astronomy and Radio Physics A. Krüger, 2012-12-06 1. 1. Short History of Solar Radio Astronomy Since its birth in the forties of our century, solar radio astronomy has grown into an extensive scientific branch comprising a number of quite different topics covering technical sciences, astrophysics, plasma physics, solar-terrestrial physics, and other disciplines. Historically, the story of radio astronomy goes back to the times of James Clerk Maxwell, whose well known phenomenological electromagnetic field equations have become the basis of present-time radio physics. As a direct consequence of these equations, Maxwell was able to prognosticate the existence of radio waves which fifteen years later were experimentally detected by the famous work of Heinrich Hertz (1887/88). However, all attempts to detect radio waves from cosmic objects failed until 1932, which was mainly due to the early stage of development of

receiving techniques and the as yet missing knowledge of the existence of a screening ionosphere (which was detected in 1925). Therefore, famous inventors like Thomas Edison and A. E. Kennelly, as well as Sir Oliver Lodge, were unsuccessful in receiving any radio emission from the Sun or other extraterrestrial sources. Another hindering point was that nobody could a priori expect that solar radio emission should have something to do with solar activity so that unfortunately by chance some experiments were carried out just at periods of low solar activity. This was also why Karl Guthe Jansky at the birth of radio astronomy detected galactic radio waves but no emission from the Sun.

Coronal Physics from Radio and Space Observations Gerard Trottet,1997-04-17 This volume provides an overview of our current understanding of the physics related to: coronal structures and coronal heating; large-scale coronal shock waves and coronal mass ejections; particle beams in the solar corona and in the interplanetary medium; and explosive energy-release phenomena and particle acceleration. The different articles give a well-balanced presentation of relevant observations based upon various techniques, models and theories, providing a global view of these phenomena and of the underlying physics. In-situ measurements of particles and waves with ULYSSES and WIND and spectral and imaging data from SOHO and YOHKOH provide an unprecedented richness of relevant data. For their better understanding, radio observations - also included in this book - play a key role.

The Solar Corona Leon Golub, Jay M. Pasachoff, 2010 Second edition graduate level textbook giving an up-to-date treatment of our understanding of the solar corona.

Proceedings - Institution of Radio and Electronics Engineers Australia Institution of

Radio and Electronics Engineers, Australia, 1971

Initial Results from the Nobeyama Radioheliograph Nobeyama Uchū Denpa Kansokujo (Japan), Nihon Tenmon Gakkai, 1994

New Millennium Solar Physics Markus J. Aschwanden, 2019-05-22 This is a followon book to the introductory textbook Physics of the Solar Corona previously published in 2004 by the same author, which provided a systematic introduction and covered mostly scientific results from the pre-2000 era. Using a similar structure as the previous book the second volume provides a seamless continuation of numerous novel research results in solar physics that emerged in the new millennium (after 2000) from the new solar missions of RHESSI, STEREO, Hinode, CORONAS, and the Solar Dynamics Observatory (SDO) during the era of 2000-2018. The new solar space missions are characterized by unprecedented high-resolution imaging, time resolution, spectral capabilities, stereoscopy and tomography, which reveal the intricate dynamics of magneto-hydrodynamic processes in the solar corona down to scales of 100 km. The enormous amount of data streaming down from SDO in Terabytes per day requires advanced automated data processing methods. The book focuses exclusively on new research results after 2000, which are reviewed in a comprehensive manner, documented by over 3600 literature references, covering theory, observations, and numerical modeling of basic physical processes that are observed in high-temperature plasmas of the Sun and other astrophysical objects, such as plasma instabilities, coronal heating, magnetic reconnection processes, coronal mass ejections, plasma waves and oscillations, or particle acceleration.

Proceedings of the STIP Symposium on Solar Radio Astronomy, Interplanetary

Scintillations and Coordination with Spacecraft M. A. Shea,1982

Astronomy in India: A Historical Perspective Thanu Padmanabhan, 2014-08-20 India has a strong and ancient tradition of astronomy, which seamlessly merges with the current activities in Astronomy and Astrophysics in the country. While the younger generation of astronomers and students are reasonably familiar with the current facilities and the astronomical research, they might not have an equally good knowledge of the rich history of Indian astronomy. This particular volume, brought out as a part of the Platinum Jubilee Celebrations of Indian National Science Academy, concentrates on selected aspects of historical development of Indian astronomy in the form of six invited chapters. Two of the chapters - by Balachandra Rao and M.S. Sriram - cover ancient astronomy and the development of calculus in the ancient Kerela text Yuktibhasa. The other four chapters by B.V. Sreekantan, Siraj Hasan, Govind Swarup and Jayant Narlikar deal with the contemporary history of Indian astronomy covering space astronomy, optical astronomy, radio astronomy and developments in relativistic astrophysics. These chapters, written by experts in the field, provide an in-depth study of the subject and make this volume quite unique.

Energy Conversion and Particle Acceleration in the Solar Corona Karl-Ludwig Klein,2008-01-11 The conversion of energy generated in the Sun's interior creates its hot corona and a wealth of dynamical phenomena such as flares and mass ejections. Based on recent significant progress in understanding magnetic reconnection and a wealth of new observations of energetic particle signatures from the sun, the present volume reviews the current theoretical and experimental status in the field. Paying attention to both the

details and the broader picture, this book addresses both the experienced researcher as well as non-specialist researchers from related areas and postgraduate students.

Physics of the Invisible Sun Ashok Ambastha, 2020-03-27 Physics of the invisible Sun: Instrumentation, Observations, and Inferences provides a new updated perspectives of the dramatic developments in solar physics mainly after the advent of the space era. It focusses on the instrumentation exploiting the invisible windows of the electromagnetic spectrum for observing the outer, fainter layers of the Sun. It emphasizes on the several technical and observational challenges and proceeds to discuss the discoveries related to energetic phenomena occurring in the transition region and corona. The book begins with giving a brief glimpse of the historical developments during the pre-, and post-telescopic periods of visible and spectroscopic techniques, ground-based optical and radio observing sites. Various types of telescopes and back-end instrumentation are presented based on photometry, spectroscopy, and polarimetry using the Zeeman and Hanle effects for measurement of magnetic fields, and Doppler effect for radial velocity measurements. The book discusses theoretical and observational inferences based on detection of solar neutrinos, and helioseismology as the probes of the hidden solar interior, and tests of solar standard models. The characteristic properties and observational signatures of global solar p- and g-oscillations modes, developments in local helioseismology and asteroseismology are discussed. The role of the solar magnetic field and differential rotation in the activity and magnetic cycles, prediction methodologies, and dynamo models are described. Observing the Sun in IR at the longer, and the UV, EUV, XUV, X-rays, and gamma-rays at the shorter wavelengths are covered in detail. Observational challenges at

each of these wavelengths are presented followed by the instrumentation for detection and imaging that have resulted in enhancing the understanding of various solar transient phenomena, such as, flares and CMEs. The outer most corona is described as a dynamic, expanding component of the Sun from the theoretical and observational perspectives of the solar wind. It then discusses the topics of the Interplanetary magnetic field, slow and fast solar wind, interaction with magnetised and non-magnetised objects of the solar system, the space weather and the physics of the heliosphere. The chapter on the future directions in solar physics presents a brief overview of the new major facilities in various observing windows, and the future possibilities of observing the Sun from ground and vantage locations in space. Features: Systematic overview of the developments in instrumentation, observational challenges and inferences derived from ground-based and space-borne solar projects. Advances in the understanding about the solar interior from neutrinos and helioseismology. Recent research results and future directions from groundand space-based observations. This book may serve as a reference book for scientific researchers interested in multi-wavelength instrumentation and observational aspects of solar physics. It may also be used as a textbook for a graduate-level course.

Illustrated Glossary for Solar and Solar-Terrestrial Physics A. Bruzek, C.J. Durrant, 2012-12-06 At the XV. General Assembly of the International Astronomical Union in Sydney 1973, Commission 10 for Solar Activity requested the incoming Organising Committee to establish a small group to recommend a standard nomenclature for solar features and to prepare an illustrated text which would clear the jungle of terms for the benefit of solar physicists as well as of theoreticians and research workers in related

fields. The challenge was taken up by the president of Commission 10, Prof. K. O. Kiepenheuer, and his persuasive advocacy has led eventually to the present book. In the course of the work, the declared aim but not the basic purpose was revised. Rather than prepare a list of standard terms, we have preferred to collect together all the terms that appear in current English-language literature. Synonyms and partially overlapping terms are all recorded for the most part without prejudice. Each has been defined as exactly as possible with the hope that in the future they may be used and understood without ambiguity. It would be a step on the road to standardisation if these terms were not reused for new phenomena. New observations and new theories will lead to reappraisals and redefinitions so the Glossary is intended more as a guide to the present situation than as a rule-book.

Solar Radio Storms A. O. Benz, Paolo Zlobec, 1983

Radio Continua During Solar Flares Arnold O. Benz,2012-12-06 Continuum radio emission and fine structure (in particular millisecond spikes) have recently raised interest as diagnostic tools for the interpretation of energy release and particle acceleration in flares. In the circles of the European solar radio astronomers, loosely organized in CESRA, the idea of a workshop came up intended for active observers of the impulsive phase of flares in radio and associated emissions. The scientific organizing committee included A.D. Benz (chairman), A. Magun, M. Pick, G. Trottet, and P. Zlobec. The workshop was held on May 27-31, 1985 in the castle of Duino near Trieste, Italy. The meeting intended to find a common terminology, to compare radio observations with measurements in other emissions and to confront observations with theoretical concepts. We have achieved a

representative summary on the current status of the field and a clear perspective for the next cycle. This volume contains the reviews and a selection of contributions and extended abstracts of papers presented at the workshop. I wish to thank the local organizers, in particular A. Abrami, M. Comari, F. Depolli, L. Fornasari, M. Messerotti (chairman), M. Nonino, and P. Zlobec. Financial support was graciously provided by the Italian Research Council (CNR). Most of all, however, I would like to express my thankfulness to our host, His Highness Prince Raimondo della Torre e Tasso, for his invaluable hospitality. We are deeply sorry to hear of his passing in the meantime. To his memory these proceedings are dedicated.

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